# UPPER MANTLE STRUCTURE DEDUCED FROM SEISMIC RECORDS ACQUIRED DURING PROJECT EDZOE IN SOUTHERN SASKATCHEWAN AND WESTERN MANITOBA BETWEEN DISTANCES OF ABOUT 790 KILOMETERS AND 1285 KILOMETERS 

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#### Abstract

In August 1969, the Seismology Division of the Dominion Observatory detonated a series of chemical explosions in Greenbush Lake, British Columbia; the project is known as "Project Edzoe". A total of twenty explosions were attempted in 180 feet of water. The seismic field crew from the Department of Earth Sciences, University of Manitoba, obtained eight seismic records along an east-west profile in southern Saskatchewan and Manitoba; recording distances were in the range 790 to 1285 kilometers.

Signal frequencies on the records were less than 7 Hz ; noise frequencies were generally above 7 Hz . Analog playbacks increased the signal to noise ratio by about 68 percent; digital filters offered no improvement over analog playbacks.

An upper mantle velocity structure consisting of a linear velocity-depth gradient, below the base of the crust, accounts for first arrival times. However, uncertainty of crustal structure beneath the shot point and recording sites produces uncertainty in the velocity at the base of the crust and the velocity gradient immediately below it. A second arrival, following the first within about one second, can be explained by a rapid increase in velocity gradient


occurring between depths of about 120 and 150 kilometers. Evidence is given for the existence of a very low gradient following the rapid increase.

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## CHAPTER I

## INTRODUCTION

In August 1969, the Seismology Division of the Dominion Observatory detonated a series of chemical explosions in Greenbush Lake, British Columbia. The purpose of the explosions was to assist Canadian and U.S. universities and government agencies to carry out crustal and upper mantle investigations. A total of 20 explosions were attempted in 180 feet of water. A single component instrument was maintained at Lumby, 88 kilometers from the shot point. The project is called "Project Edzoe".

The seismic field crew from the Department of Earth Sciences, University of Manitoba, successfully obtained eight seismic records along an east-west profile in southern Saskatchewan and Manitoba. The recording equipment, which includes the Texas Instruments Incorporated VLF-2 refraction system, is described by Hajnal (1970). The recording stations have been given the names $S 1$ to 58 . Figure 1 shows the locations of the shot point and recording sites; Fig. 2 shows the recording site geometry. Locations and distances of the recording sites, and times, charge weights and amplitudes at Lumby of the corresponding shots are given in Table I. Locations and distances of recording sites are given for


Fig. 2. Recording site geometry.


| N | ${ }_{\sim}^{*}$ | $\infty$ | ${ }_{\infty}^{\infty}$ | $\cdots$ | の | $\cdots$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |



geophone (channel) 1. The shot corresponding to record Sl was a partial misfire due to improper priming. Distances were determined from Fortran program "l Origin Many Locations" kindly provided by the Department of Physics, University of Alberta. The program computes distances on the basis of the 1924 international constants for the reference ellipsoid (International Dictionary of Geophysics, 1967).

## CHAPTER II

## SEISMIC FILTERS

The purpose of this chapter is to give bases of description of seismic filters and to describe existing filtering techniques on these bases.

## Bases of Description of Seismic Filters

1. The purpose of any seismic filter is to extract signal from a seismic record consisting of signal plus noise. The observer defines those seismic events which are signal and those which are noise. Once signal and noise are defined, a seismic filter is accordingly defined. Thus, filtering techniques are described in terms of the definition of signal and noise.
2. The general filtering method is also a basis of description. General filtering methods include physical, mathematical or digital, and electronic methods.
3. The specific filtering method applied is the final basis of description. This part of the description includes exact mathematical, electronic or physical details of the technique.

## Glossary of Seismic Filtering Techniques

Following is a description of several, but by no means all, filtering techniques which are presently employed in seismology.

Array wavelength filter: a physical deconvolution filter. Seismic events of a specified wavelength are rejected by means of suitable shotpoint-detector geometry. Usually Rayleigh and Love waves are considered to be noise. Holzman (1963) describes how Chebyshev polynomials may furnish optimum shot-detector geometries for given problems. Roden (1965) applies wavelength filtering to teleseisms. Deconvolution filter: a filter which performs the inverse process of any of the filtering processes resulting from the passage of seismic energy through the earth. Deconvolution is equivalent to the commonly used term 'inverse convolution'. Rice (1962) discusses a mathematical approach to inverse convolution filters.

Frequency band-pass filter: a filter which accepts those seismic events within a certain frequency range and rejects those events which are outside of this range. This type of filter is effective when there is a marked separation between signal and noise frequencies. Frequency filters may be electrical or digital.

Laser beam filtering: is effective in removing both coherent noise and incoherent noise from a seismic variable density record. Velocity and frequency filtering by means of laser
beam are discussed by Dobrin et al. (1965, 1967). Motion product filter: a physical filter which combines voltages of a three component seismometer in order to suppress random noises arriving from all directions. White (1964) describes such a filter.

Multichannel filter: a filter which acts on more than one trace of a seismic record. Any multichannel filter inherently uses redundancy as a noise reducing mechanism. Multiple reflection deconvolution filter: (also called "Ghost Elimination Filter") a deconvolution filter which separates primary reflections (signal) from multiple reflection (noise). Lindsey (1960) discusses the realization of such a filter by means of an analog feedback system. Goupillaud (1961) uses a direct approach to filtering multiples. Hammond (1962) describes a physical ghost elimination filter. Silverman et al. (1963) approach the problem of multiples by means of "Murac", an analog computer. Schneider et al. (1965) combine multichannel digital filtering with stacking to remove primary reflections from multiples plus noise. Anstey et al. (1966) show the effectiveness of sectional auto-correlograms and sectional retrocorrelograms in separating primary and multiple reflections. Non-Linear filter: a filter designed for non-stationary time series input. Robinson (1967) presents non-linear filter theory. Clarke (1968) describes time varying deconvolution filters.

Optimum filter: a filter which is designed on the basis of some optimality condition; because of complexities generally encountered as a result of the optimality conditions, optimum filters are usually digital.

Predictive filter: a filter which removes random events from a seismic record by predicting future values of a given stationary stochastic process. Usually predictive filters are Wiener filters; the prediction operator is calculated such that the predicted output is as close as possible (in the least squares sense) to a particular desired output: Predictive filters are described in detail by Robinson(1967). Predictive deconvolution filter: a predictive filter which removes undesirable seismic energy responses caused by the earth. Discussion is given by Robinson (1967). Recursive filter: a filter which produces output which is a function of both input and past output values. The term "recursive" is used when the filter is digital; an electrical recursive filter is called a feedback filter. Meyerhoff (1966) describes a combination stacking and optimum feedback system. Shanks (1967) gives a general discussion of recursive filters. Usually recursive filters are computationally efficient.

Stacking fiZter: a filter which removes random noise by means of simple addition of several seismic traces. The terms "multiple coverage" and "common depth (reflection) point" are associated with the stacking technique; channels representing common reflection points are stacked to remove
random noise. Stacking filters are more effective than frequency filters when there is an overlap in signal and noise frequencies, however, they are not designed for the removal of coherent noise. Stacking filters are discussed by Mayne (1962) and Galbraith et al. (1968).

Velocity filter: a deconvolution filter which accepts all seismic events within a specified apparent velocity band and rejects seismic events outside of this band. Thus noise of frequency and wavelength, which fall in the signal frequency and wavelength range, may often be removed on the basis of apparent velocity separation. A fan filter is a velocity filter which passes events that have apparent velocities which fall within a certain fan-shaped region in the frequency-wavenumber plane. The "pie-slice" filter described by Embree et al. (1963) is another example of a velocity filter.

Wiener fizter: an optimum filter. The optimality condition is that the actual output be as close as possible to some specified desired output (in the least squares sense). Wiener filtering is described by Wiener (1949) and Robinson (1967).

Names which have been given to filtering techniques in the literature have been derived on the basis of either broad or specific characteristics of the technique. As a result of this, any specific filter may have a combination of titles. An example of this is the "optimum multichannel
velocity deconvolution filter" described by Sengbush et al. (1968).

## CHAPTER III

## FILTERING TECHNIQUES EMPLOYED

Signal frequencies are below 7 Hz and noise frequencies are generally above 7 Hz on the records from the present experiment. Digital and analog filtering were used in an attempt to remove white noise from the records; it was found that digital filtering offered no advantages over simple analog playback filtering.

## Digital Filtering

Analog to digital conversion was carried out by means of the Radiation Inc. A/D converter. All twelve channels were digitized for each record. The digitizing interval used was 1.71 milliseconds; hence, the aliasing frequency was about 300 Hz , which is well above signal frequencies. The Radiation converter is described in detail by Hajnal (1970). All digital processing was done on the IBM $360 / 65$ at the Department of Computer Science, University of Manitoba. Digital seismic data was plotted by means of the Calcomp 750/563 plotting system.

Multichannel Digital Prediction Filtering
Multichannel digital prediction filtering was attempted
by means of Fortan computer programs written by Burgess (1969). The programs are based on the mathematical theory of prediction described by Robinson (1967). Basically, the program package written by Burgess consists of: "predict l", a program which computes a general multichannel least squares Wiener filter; and "mftconv", a subroutine which performs multichannel convolution of this filter with segmented input. The eight vertical traces on a few sample seismic records were processed by this technique. Results were unsuccessful; the normalized prediction error of realizable length optimum filters was approximately 0.7. A large prediction error in this case could be attributed to the fact that there were not enough traces (only eight) to comprise a sufficient multichannel stationary random process.

In any event, this multichannel prediction technique is designed for the prediction of first arrivals only.

## Low-pass Digital Filtering

A digital seismic filter Fortran program package has been written for the Department of Earth Sciences, University of Manitoba, by Hajnal (1970). Programs within this package used were:

Bandpass: a program which computes weighting coefficients of a bandpass filter and determines the frequency response of the computed filter. Convolv: filters the seismic data with a set of weighting coefficients calculated by "bandpass".

Plotmod: prepares seismic data for plotting.
The bandpass filter computed by "bandpass" has weighting coefficients, $b_{t}$, defined by

$$
\begin{gather*}
b_{t}=\frac{1}{t}\left\{\sin \left[2 \pi\left(h+f_{0}\right) t\right]-\sin \left[2 \pi\left(h-f_{0}\right) t\right]\right\}\left(\left.1-\frac{t}{n} \right\rvert\,\right) \\
-n<t<+n \tag{3•1}
\end{gather*}
$$

where,

$$
\begin{aligned}
f_{o} & =\text { center frequency of ideal bandpass filter. } \\
h & =\text { half-width of ideal bandpass filter. } \\
(l & \left.-\frac{t}{n}\right)=\text { Fejer weighting factor. } \\
t & =\text { time. }
\end{aligned}
$$

Figure 3 shows the frequency response curves of digital low-pass filters $F_{1}, F_{2}, F_{3}, F_{4}$, described in Table II. An increase in length from 100 to 200 results in a marked improvement in frequency response for both (0-5) Hz and (0-10) Hz filters.

Figure 3 suggests that $F_{4}$ would be effective in increasing the signal to noise ratio of the seismic data. Figures 4 and 5 show record 55 unfiltered and filtered with $\mathrm{F}_{4}$ respectively. The signal to noise ratio of the unfiltered record is approximately 1.6 ; the signal to noise ratio of the $\mathrm{F}_{4}$ filtered record is approximately 2.7. Thus, the signal to noise ratio is increased by about 68 percent.

There is a certain ambiguity in the definition of the signal frequency band. Even though each seismic phase consists of frequencies of 7 Hz or less, the superposition of two or more seismic phases contains small wavelets of



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frequency as high as 12 Hz . Small interference wavelets are valuable in determining the onset of seismic phases. In view of this fact, it was necessary to have a compromising low-pass filter which would remove a sufficient amount of noise and leave a sufficient amplitude of wavelets such that the net effect would be the production of readable seismic records. Filter $F_{4}$ is such a compromise.

Despite the fact that good digitally filtered records can be obtained with existing programs, "convolv" has disadvantages. The seismic data are stored on tape in the form of 1202 two-byte words. The first two words of a block identify the record and block numbers; the other 1200 words consist of 100 samples from each of the 12 seismic channels (0.171 seconds of seismic information per channel). During the convolution process, five blocks of data are read into core at a time (this is about 0.885 seconds of seismic information per channel). Thus, when it is desirable to process large amounts of data ( 25 seconds or more), "convolv" becomes input-output bound. Furthermore, "convolv" performs convolution in the time domain, and this is a slow process. In present form, "convolv" uses about 75 minutes CPU time to process 25 seconds of seismic data for twelve channels, a digitizing rate of 1.710 milliseconds, and a filter of length 200. Thus, for large amounts of data, "convolv" is also CPU bound. The efficiency of "convolv" could be increased by,

1. decreasing the digitizing rate,
2. decreasing filter length,
3. performing convolution in the frequency domain with the fast Fourier transform method described by Robinson (1967).

## Analog Filtering

Figure 6 shows the frequency response characteristics of the VLF-2 system. The curve for 8 Hz is very similar to the curve which describes filter $\mathrm{F}_{4}$ in Fig 3. Figure lle is an ( $0-8 \mathrm{~Hz}$ ) analog playback of record $\mathrm{S5}$; this is to be compared with record S5 filtered with digital low-pass filter $\mathrm{F}_{4}$ (Fig. 5). Very good results were obtained with (0-8 Hz) analog playbacks; thus, it was not necessary to use digital filtering. Figures lla to llh are the ( $0-8 \mathrm{~Hz}$ ) analog playbacks of records Sl to s 8.

For long range refraction experiments, signal frequencies are very low; the VLF-2 system is designed for such experiments. Digital bandpass filtering would be an improvement over analog playback filtering for studies such as near vertical reflection experiments, in which signal frequencies are higher. In cases such as these, severe limitations are placed on the analog playback system. Hajnal (1970) shows vast improvements by digital filtering techniques on near vertical reflection records.

Fig. 6. Frequency response of the VLF-2 refraction system.

## CHAPTER IV

## INTERPRETATION

An interpretation of the first two arrivals $\mathrm{P}_{1}$ and $\mathrm{P}_{2}$, is given in this chapter. Both arrival were found to be the result of rays which penetrate the upper mantle. Unfortunately, no detailed information about the crustal structure under the shot point and recording sites has been published. Uncertainty of crustal structure resulted in uncertainty in mantle structure deduced from mantle arrivals. $P_{1}$ and $P_{2}$ are shown on the records in Figs. lla to llh. Determination of Upper Mantle Velocity
from First Arrival Times

For the distance range of this experiment, rays which penetrate the upper mantle emerge as first arrivals. To account for observed first arrival times, upper mantle linear velocity-depth functions of the form given by equation $4 \cdot 1$ were considered.

$$
\begin{gathered}
v_{p}=V_{m}, z=z_{m} \\
v_{p}=v_{m}+M\left(z-z_{m}\right), Z \geq z_{m}, M \geq 0 \quad \ldots(4 \cdot 1)
\end{gathered}
$$

where

$$
\begin{aligned}
V_{p}= & P \text { wave velocity } \\
V_{m}= & \text { velocity at the base of the crust } \\
\mathrm{Z}_{\mathrm{m}}= & \text { depth to the base of the crust } \\
M= & \text { linear velocity-depth gradient in the upper } \\
& \text { mantle (seconds }{ }^{-1} \text { ) }
\end{aligned}
$$

Ranges of acceptable values of $V_{m}$, the velocity at the base of the crust, and $M$, the upper mantle gradient, depend upon not only the actual first arrival times, but also the choice of crustal structure and required accuracy of travel times.

## Uncertainty in Crustal Structure

Despite the fact that there is no detailed crustal information, it is still possible to restrict the crustal structure to certain ranges. McConnell and McTaggert-Cowan (1963) have calculated the mean crustal velocity and depth to Moho for shields and stable interior platforms (Table III). A range of crustal velocities, $6.34 \pm 0.27 \mathrm{~km} / \mathrm{sec}$, and crustal thicknesses, $41.06 \pm 7.78$ kilometers, produces a range in crustal delay times associated with rays which travel through the upper mantle. Acceptable values of $V_{m}$ and $M$ were determined for the following crustal models:

Crustal Model A; velocity $=$ mean $=6.34 \mathrm{~km} / \mathrm{sec}$ crustal thickness $=$ mean $=41.06 \mathrm{~km}$

This model produces an average crustal delay time.

## Table III

Mean Crustal Structure for Shields and Stable Interior Platforms (McConnell and McTaggart-Cowan, 1963)

|  | Crustal Velocity <br> $(\mathrm{km} / \mathrm{sec})$ | Crustal Thickness <br> $\quad$(km) |
| :--- | :---: | :---: |
| Mean | 6.34 | 41.06 |
| Standard | 0.27 | 7.78 |
| Deviation |  |  |

$$
\begin{aligned}
\text { Crustal Model B; velocity } & =\text { mean }- \text { standard deviation } \\
& =6.07 \mathrm{~km} / \mathrm{sec} \\
\text { crustal thickness } & =\text { mean }+ \text { standard deviation } \\
& =48.84 \mathrm{~km}
\end{aligned}
$$

This model produces a maximum delay time.

Crustal Model C; velocity $=$ mean + standard deviation $=6.61 \mathrm{~km} / \mathrm{sec}$
crustal thickness $=$ mean - standard deviation $=33.28 \mathrm{~km}$

This model results in a minimum delay time.

Determination of $V_{m}$ and $M$

Allowable ranges of $V_{m}$ and $M$ for a given crustal model and a specified accuracy of arrival times were determined by examining the normalized root mean square error function, $\mathrm{E}_{1}$, defined by:

$$
E_{1}=\sqrt{\frac{\sum_{i=1}^{N}\left(T_{i}-t_{i}\right)^{2}}{N}}
$$

where
$T_{i}=$ theoretical first arrival time at the $i^{\text {th }}$ station
$t_{i}=$ observed first arrival time at the $i^{\text {th }}$ station $\mathrm{N}=$ number of stations at which the first arrival is observed $=8$

For a chosen crustal model, $\mathrm{E}_{1}$ is clearly a function of $V_{m}$ and $M$ on $l y$, since each $t_{i}$ is a constant. $A$ minimum in the
function $E_{1}\left(V_{m}, M\right)$ occurs when

$$
\frac{\partial E_{1}}{\partial V_{m}}=\frac{\partial E_{1}}{\partial M}=0
$$

and corresponds to optimum values (in the least squares sense) of $V_{m}$ and $M . \quad E_{1}\left(V_{m}, M\right)$ could possibly be written as an explicit function of $V_{m}$ and $M$; optimum values of $V_{m}$ and $M$ could then be determined by solving equation $4 \cdot 2$. However, since the parametric equations relating time and distance for a spherical geometry and linear velocity-depth functions are very complicated, calculations have been performed using the IBM $360 / 65$ computer.

Table IV shows calculated values of $E_{1}\left(s e c o n d s^{-1}\right)$ for various values of $\mathrm{V}_{\mathrm{m}}$ and M , assuming the average crustal model $A$; the table shows trends in the function $E_{1}$. For each of the values of $V_{m}$ between 7.90 and 8.10 , there is a value of $M$ between 0.0005 and 0.007 which corresponds to a minimum in $E_{1}$. For $V_{m}=8.15$ and $V_{m}=8.20$, the table suggests a minimum in $E_{1}$ will be found for $M$ less than 0.0005. For each of the values of $M$ between 0.0005 and 0.006 there is a value of $\mathrm{V}_{\mathrm{m}}$ between 7.90 and 8.20 which corresponds to a minimum in $E_{1}$. For $M=0.007$ a minimum will occur when $V_{m}$ is less than 7.90.

The scanning grid of $\left(V_{m}, M\right)$ values in Table IV is neither fine enough nor extensive enough to determine acceptable solutions for a given error; the purpose of the table is to show the general nature of the error function $E_{1}$.
-
Error Function,

$$
\begin{array}{cc}
\mathrm{V}_{\mathrm{m}}(\mathrm{~km} / \mathrm{sec}) & \\
\mathrm{M}^{\left(\mathrm{sec}^{-1}\right)} & 7.90 \\
0.0005 & 3.30 \\
0.001 & 3.15 \\
0.002 & 2.75 \\
0.003 & 2.22 \\
0.004 & 1.60 \\
0.005 & 1.03 \\
0.006 & 1.00 \\
0.007 & 1.74
\end{array}
$$

The error function has been calculated with a grid spacing of $\Delta V_{m}=0.02 \mathrm{~km} / \mathrm{sec}$ and $\Delta M=0.0001$ seconds $^{-1}$ for each of crustal models A, B and C.

Table $V$ shows the results for crustal model A. For each value of $V_{m}$, values of $E_{1}$ are given for the range of $M$ which shows $E_{1}$ passing through a minimum. Each value of $V_{m}$ has a minimum value of $E_{1}$ which corresponds to an optimum M. The set of minimum values of $E_{1}$ also has a minimum; this is shown in Table VI. It should be noted that the minimum value of $E_{1}$ for $V_{m}=8.14 \mathrm{~km} / \mathrm{sec}, V_{m}=8.16$ $\mathrm{km} / \mathrm{sec}$, and $V_{m}=8.18 \mathrm{~km} / \mathrm{sec}$ has been taken as the value corresponding to $M=0$, since negative gradients have not been considered.

In accordance with Table $V I$, the minimum value of $\mathrm{E}_{1}$ for any $V_{m}$ less than $7.98 \mathrm{~km} / \mathrm{sec}$ must be greater than 0.57 seconds and the minimum value of $E_{1}$ for any $V_{m}$ greater than $8.18 \mathrm{~km} / \mathrm{sec}$ must be greater than 0.65 seconds. Table $V$ furnishes acceptable values of $V_{m}$ and $M$ for $a$ given error in observed arrival times (assuming crustal model A).

For each crustal model acceptable values of $V_{m}$ of the form,

$$
\mathrm{v}_{\mathrm{a}} \leq \mathrm{v}_{\mathrm{m}} \leq \mathrm{v}_{\mathrm{b}}
$$

were found for first arrival time accuracies of 0.50 seconds and 0.30 seconds. For each value of $V_{m}$ between $V_{a}$ and $V_{b}$
Table $V$
Error Function $E_{1}\left(V_{m}, M\right)$ for Crustal Model $A$
$\begin{array}{cc}\mathrm{V}_{\mathrm{m}}=8.08 & \mathrm{~km} / \mathrm{sec} \\ \mathrm{M} & \mathrm{E}_{1} \\ \left(\mathrm{sec}^{-1}\right) & (\mathrm{sec}) \\ 0.0013 & .51 \\ 0.0014 & .48 \\ 0.0015 & .44 \\ 0.0016 & .41 \\ 0.0017 & .38 \\ 0.0018 & .34 \\ 0.0019 & .31 \\ 0.0020 & .29 \\ 0.0021 & .26 \\ 0.0022 & .25 \\ 0.0023 & .24 \\ 0.0024 & .25 \\ 0.0025 & .27 \\ 0.0026 & .29 \\ 0.0027 & .33 \\ 0.0028 & .37 \\ 0.0029 & .41 \\ 0.0030 & .46 \\ 0.0031 & .52\end{array}$


$$
\begin{array}{cc}
\mathrm{V}_{\mathrm{m}}=8.16 & \mathrm{~km} / \mathrm{sec} \\
\left.\mathrm{M}^{-1}\right) & \mathrm{E}_{1} \\
\left(\mathrm{sec}^{-1}\right) & (\mathrm{sec}) \\
0 & .41 \\
0.0001 & .43 \\
0.0002 & .44 \\
0.0003 & .45 \\
0.0004 & .47 \\
0.0005 & .49 \\
0.0006 & .51
\end{array}
$$

$$
\begin{array}{cc}
\mathrm{V}_{\mathrm{m}}=8.18 & \mathrm{~km} / \mathrm{sec} \\
\mathrm{M} & \mathrm{E}_{1} \\
\left(\mathrm{sec}^{-1}\right) & (\mathrm{sec}) \\
0 & .65 \\
0.0001 & .66 \\
0.0002 & .68
\end{array}
$$

(pənutquos) $\Delta$ əтqед

$$
\begin{array}{cc}
\mathrm{V}_{\mathrm{m}}=8.10 & \mathrm{~km} / \mathrm{sec} \\
\mathrm{M} & \mathrm{E}_{1} \\
\left(\mathrm{sec}^{-1}\right) & (\mathrm{sec}) \\
0.0004 & .51 \\
0.0005 & .49 \\
0.0006 & .46 \\
0.0007 & .44 \\
0.0008 & .41 \\
0.0009 & .39 \\
0.0010 & .36 \\
0.0011 & .34 \\
0.0012 & .31 \\
0.0013 & .29 \\
0.0014 & .26 \\
0.0015 & .25 \\
0.0016 & .23 \\
0.0017 & .22 \\
0.0018 & .22 \\
0.0019 & .23 \\
0.0020 & .25 \\
0.0021 & .28 \\
0.0022 & .31 \\
0.0023 & .35 \\
0.0024 & .39 \\
0.0025 & .43 \\
0.0026 & .48 \\
0.0027 & .53 \\
0.43
\end{array}
$$

$$
\begin{aligned}
& 8.00 \\
& 0.50
\end{aligned}
$$

$$
\begin{aligned}
& 8.12 \\
& 0.23
\end{aligned}
$$

$$
\begin{array}{ll}
\infty & n \\
\stackrel{\infty}{r} & \vdots \\
\stackrel{0}{\circ} & 0
\end{array}
$$

there is an acceptable range of $M$ of the form,

$$
M_{a} \leq M \leq M_{b}, M_{a} \geq 0
$$

Table VII lists $V_{a}$ and $V_{b}$ for the three crustal models. Table VIII gives $M_{a}$ and $M_{b}$ corresponding to acceptable values of $\mathrm{V}_{\mathrm{m}}$. For a given crustal model, acceptable values of $\mathrm{V}_{\mathrm{m}}$ which are lower have higher values of $M$. Crustal model A represents an average crustal delay time; accordingly, solution values of $V_{m}$ and $M$ are intermediate. Crustal model $B$ represents a maximum delay time; high values of $V_{m}$ and low values of $M$ are required for a solution. Crustal model $C$ represents a minimum delay time; low values of $V_{m}$ and high values of $M$ are required.

The effect of crustal structure upon the solution of upper mantle velocity is more easily seen by comparing the best solutions of $V_{m}$ and $M$ for each crustal structure. Solutions $A, B$ and $C$ are the best solutions of $V_{m}$ and $M$ assuming crustal structures $A, B$ and $C$ respectively;

Solution $A: \quad V_{m}=8.10 \mathrm{~km} / \mathrm{sec}, \quad \mathrm{M}=0.0017 \mathrm{sec}^{-1}$,

$$
\mathrm{E}_{1}=0.22 \mathrm{sec}
$$

Solution $B: \quad V_{m}=8.32 \mathrm{~km} / \mathrm{sec}, \quad M=0 \mathrm{sec}^{-1}$,

$$
\mathrm{E}_{1}=0.37 \mathrm{sec}
$$

Solution $C: \quad V_{m}=7.88 \mathrm{~km} / \mathrm{sec}, \quad \mathrm{M}=0.0036 \mathrm{sec}^{-1}$,

$$
\mathrm{E}_{1}=0.27 \mathrm{sec}
$$

Figure 7 shows solutions A, B and C graphically. Table IX lists theoretical and observed first arrival times, and

Table VII

Solutions of $\mathrm{V}_{\mathrm{m}}, \mathrm{V}_{\mathrm{a}} \leq \mathrm{V}_{\mathrm{m}} \leq \mathrm{V}_{\mathrm{b}}$
for Crustal Models $A, B$, and $C$

| Crustal <br> Model | Accuracy of <br> First Arrival <br> Times (sec) | $V_{a}$ <br> $(\mathrm{~km} / \mathrm{sec})$ | $\mathrm{V}_{\mathrm{b}}$ <br> $(\mathrm{km} / \mathrm{sec})$ |
| :---: | :---: | :---: | :---: |
| A | .50 | 8.00 | 8.16 |
|  | .30 | 8.06 | 8.14 |
| B | .50 | 8.28 | 8.34 |
|  | .30 | - | No Solutions |
| C | .50 | 7.80 | - |
|  | .30 | 7.86 | 7.96 |

Table VIII

Solutions of $M, M_{a} \leq M \leq M_{b}$, for First Arrival Time Accuracies of 0.5 sec and 0.3 sec

| $\begin{gathered} \mathrm{V}_{\mathrm{m}} \\ (\mathrm{~km} / \mathrm{sec}) \end{gathered}$ | Crustal Model A |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
|  | 0.5 sec Accuracy |  | 0.3 sec Accuracy |  |
|  | $\begin{gathered} M_{a} \\ \left(\sec ^{-1}\right) \end{gathered}$ | $\begin{gathered} M_{b} \\ \left(\sec ^{-1}\right) \end{gathered}$ | $\begin{gathered} M_{a} \\ \left(\sec ^{-1}\right) \end{gathered}$ | $\begin{gathered} M_{b} \\ \left(\sec ^{-1}\right) \end{gathered}$ |
| 8.00 | 0.0039 | 0.0040 | - | - |
| 8.02 | 0.0032 | 0.0039 | - | - |
| 8.04 | 0.0027 | 0.0037 | - | - |
| 8.06 | 0.0021 | 0.0034 | 0.0027 | 0.0029 |
| 8.08 | 0.0014 | 0.0030 | 0.0020 | 0.0026 |
| 8.10 | 0.0005 | 0.0026 | 0.0013 | 0.0021 |
| 8.12 | 0 | 0.0021 | 0.0004 | 0.0015 |
| 8.14 | 0 | 0.0014 | 0 | 0.0006 |
| 8.16 | 0 | 0.0005 | - | - |

Table VIII (Continued)

Crustal Model B

## 0.5 sec Accuracy

0.3 sec Accuracy

| $\mathrm{V}_{\mathrm{m}}$ <br> $(\mathrm{km} / \mathrm{sec})$ | $\mathrm{M}_{\mathrm{a}}$ <br> $\left(\mathrm{sec}^{-1}\right)$ | $\mathrm{M}_{\mathrm{b}}$ <br> $\left(\mathrm{sec}^{-1}\right)$ | $\mathrm{M}_{\mathrm{a}}$ <br> $\left(\mathrm{sec}^{-1}\right)$ | $\mathrm{M}_{\mathrm{b}}$ <br> $\left(\mathrm{sec}^{-1}\right)$ |
| :---: | :---: | :---: | :---: | :---: |
| 8.28 | 0.0010 | 0.0017 | - | - |
| 8.30 | 0 | 0.0014 | - | - |
| 8.32 | 0 | 0.0009 | - | - |
| 8.34 | 0 | 0.0002 | - | - |

Crustal Model C

| 7.80 | 0.0047 | 0.0052 | - | - |
| :--- | :---: | :---: | :---: | :---: |
| 7.82 | 0.0043 | 0.0050 | - | - |
| 7.84 | 0.0039 | 0.0047 | - | - |
| 7.86 | 0.0035 | 0.0044 | 0.0039 | 0.0041 |
| 7.88 | 0.0031 | 0.0041 | 0.0035 | 0.0038 |
| 7.90 | 0.0027 | 0.0038 | 0.0031 | 0.0033 |
| 7.92 | 0.0022 | 0.0033 | - | - |
| 7.94 | 0.0018 | 0.0028 | - | - |
| 7.96 | 0.0014 | 0.0022 | - | - |



Fig. 7. Best solutions for crustal models A, B, and C.
Table IX

depths of penetration of rays which emerge at the recording sites for solutions $A, B$ and $C$. The range in depth at which the rays bottom depends drastically upon the choice of crustal structure.

The velocity at the base of the crust, and the velocity gradient within the upper mantle cannot be determined accurately on the basis of first arrival times. Uncertainty of crustal structure permits a wide range of values for $V_{m}$ and $M$, even when observed first arrival times are known to an accuracy of $\pm 0.30$ seconds (Table VIII). However, the velocity-depth functions corresponding to the best solutions for a wide range of crustal structures ( $A, B$ and $C$ ) do converge with depth as shown in Fig. 7.

## Evidence of Upper Mantle Velocity Structure

from Second Arrivals

A second event, $P_{2}$, has been picked on all of the records except Sl which is of poor quality. On record S 2 , $P_{2}$ has been taken as the third energy arrival since the arrival time of the second event does not correlate with arrival times of $\mathrm{P}_{2}$ on the other records.

The possibility that $\mathrm{P}_{2}$ is either a multiple reflection at the free surface or a PS conversion has been eliminated.

Green and Hales (1968) have reported strong multiple phases (PP, PPP and PPPP) on seismic records from Project Early Rise. They point out that, theoretically, these phases should not be visible for the distances at which they
have observed them. For a velocity of $6 \mathrm{~km} / \mathrm{sec}$ in the upper part of the crust, $P_{2}$ arrives too early to be a conventional multiple reflection.

The PS converted wave is composed of SV type motion. It results from the conversion of energy, in the form of a refraction, from the parent $P$ wave at the interface between two crustal layers of contrasting seismic velocity. Schwind et al. (1960) find various multiple PS conversions on seismic records up to about 400 kilometers. However, amplitude curves of McCamy et al. (1962) show that the ratio of converted $P S$ wave amplitude to parent $P$ wave amplitude is very small for distances of this experiment. For example, for a conversion at a boundary separated by seismic $P$ wave velocities of $6.0 \mathrm{~km} / \mathrm{sec}$ and $6.5 \mathrm{~km} / \mathrm{sec}$, the amplitude ratio is less than 0.05 at a distance of about 1000 kilometers.

## Effect of Rapid Increase in Velocity Gradient

A rapid increase in velocity gradient produces a triplication in a time-distance plot (Fig. 8). It was assumed that $P_{2}$ was part of a triplication corresponding to branch $X Y$ in Fig. 8. Observed arrival times of $P_{2}$ were explained by a velocity gradient $M_{2}$ commencing at a depth $Z_{2}$ such that $M_{2}$ was greater than the velocity gradient above $Z_{2}$. For each of solutions $A, B$, and $C$, upper mantle velocity functions of the following form were considered:


$$
\begin{aligned}
& v_{p}=V_{1}+M_{1}\left(z-Z_{1}\right), Z_{2} \geq z \geq z_{1} \\
& v_{p}=V_{1}+M_{1}\left(Z_{2}-Z_{1}\right)+M_{2}\left(z-z_{2}\right), z \geq z_{2}
\end{aligned}
$$

where

$$
\begin{aligned}
V_{p}= & P \text { wave velocity in upper mantle } \\
\left(V_{1},\right. & \left.M_{1}\right)=\text { best values (in the least squares sense) of } \\
& V_{m} \text { and } M \text { for solution } A, B \text { or } C . \\
Z= & \text { depth } \\
Z_{1}= & \text { crustal thickness or crustal model corresponding } \\
& \text { to solution } A, B \text { or } C \\
Z_{2}= & \text { depth at which the velocity gradient becomes } M_{2} \\
M_{2}= & \text { new upper mantle gradient such that } M_{2}>M_{1}
\end{aligned}
$$

$Z_{2}$ and $M_{2}$ were restricted to values for which:
i) the point $Y$, on Fig. 8 , occurs at a distance greater than 1284 kilometers (the largest distance at which $P_{2}$ is observed).
ii) the point $W$, on Fig. 8, occurs at a distance less than 831 kilometers (the smallest distance at which $P_{2}$ is observed).

For each of solutions A, B and C, a normalized root mean square error function, $E_{2}$, was calculated for increments of 10 kilometers in $Z_{2}$ and various values of $M_{2}$ :

$$
E_{2}=\sqrt{\frac{\sum_{i=2}\left(T_{i}-t_{i}\right)^{2}}{N}}
$$

where
$T_{i}=$ theoretical arrival time of $P_{2}$ at the $i$ th station $t_{i}=$ observed arrival time of $P_{2}$ at the $i t h$ station $N=$ number of stations at which $P_{2}$ is observed $=7$

The choice of solutions A, B or C (and, in turn, the choice of crustal model A, B or C) allows variation in the depth, $Z_{2}$, at which a rapid increase in velocity can take place. Following are the values of $Z_{2}$ and $M_{2}$ for which $E_{2}$ is less than about 0.5 seconds.

Calculations assuming Solution A (Crustal model A):

$$
\begin{aligned}
& \mathrm{z}_{2}=120 \text { kilometers, } \mathrm{M}_{2}>0.0155 \pm 0.0005 \mathrm{sec}^{-1} \\
& \mathrm{z}_{2}=130 \text { kilometers, } \mathrm{M}_{2}>0.0185 \pm 0.0005 \mathrm{sec}^{-1} \\
& \mathrm{z}_{2}=140 \text { kilometers, } \mathrm{M}_{2}>0.045 \pm 0.005 \mathrm{sec}^{-1}
\end{aligned}
$$

Calculations assuming Solution B (Crustal modeZ B):

$$
\mathrm{z}_{2}=120 \text { kilometers, } \mathrm{M}_{2}>0.025 \pm 0.005 \mathrm{sec}^{-1}
$$

Caloulations assuming Solution $C$ (Crustal model C):

$$
\begin{aligned}
& \mathrm{z}_{2}=140 \text { kilometers, } \mathrm{M}_{2}>0.0205 \pm 0.0005 \mathrm{sec}^{-1} \\
& \mathrm{z}_{2}=150 \text { kilometers, } \mathrm{M}_{2}>0.055 \pm 0.005 \mathrm{sec}^{-1}
\end{aligned}
$$

The best value of $E_{2}$ is about 0.4 sec; it occurs when the average crustal model A is assumed and $\mathrm{Z}_{2}=130$ kilometers, $M_{2}>0.025 \pm 0.005 \mathrm{sec}^{-1}$.
$\underline{\mathrm{P}_{2} / \mathrm{P}_{1} \text { Amplitude Ratio: }}$
For each of solutions $A, B$ and $C$, the theoretical $P_{2} / P_{1}$ amplitude ratios, based on geometric spreading, were
calculated for values of $Z_{2}$ and $M_{2}$ for which $E_{2}$ is less than about 0.5 seconds. It was found that the $P_{2} / P_{1}$ ratio does not change appreciably with either a change in solution (crustal model) or changes in ( $\mathrm{Z}_{2}, \mathrm{M}_{2}$ ) values for a given solution (crustal model). Thus, acceptable ( $\mathrm{Z}_{2}, \mathrm{M}_{2}$ ) values cannot be restricted further on the basis of amplitude ratios. However, the theoretical $P_{2} / P_{1}$ amplitude ratios do show general agreement with observed $P_{2} / P_{1}$ amplitude ratios. Figure 9 shows the theoretical $P_{2} / P_{1}$ ratios for solution $A$ ( $Z_{2}=130$ kilometers, $M_{2}=0.10 \mathrm{sec}^{-1}$ ) plotted against observed values.

Velocity changes sufficient to explain $\mathrm{P}_{2}$
It has been shown that $P_{2}$ arrival times and $P_{2} / P_{1}$ amplitude ratios can be explained if the velocity gradient suddenly increases between depths of about 120 kilometers and 150 kilometers. The value of the new gradient, $M_{2}$, is bounded below but it may tend to infinity. However, $M_{2}$ need only exist to a depth $Z_{3}$ such that $P_{2}$ will theoretically be observed at a distance of about 831 kilometers (the distance of station S2). Table $x$ gives $Z_{3}, V_{p}\left(Z_{3}\right)$, and $V_{p}\left(Z_{2}\right)$ for solutions $A, B$ and $C$, and for acceptable values of $Z_{2}$ and $M_{2}$. A linear change of velocity from $V_{p}\left(Z_{2}\right)$ to $V_{p}\left(Z_{3}\right)$ between depths of $Z_{2}$ and $Z_{3}$ is sufficient to explain the existence of the $\mathrm{P}_{2}$ phase at distances as small as 831 kilometers. For a given solution and a given depth $Z_{2}$ at which $M_{2}$ begins, clearly $Z_{3}$ is a function of $M_{2}$ only. Furthermore, $V_{p}\left(Z_{3}\right)$
Table X

| Sufficient Depth, $\mathrm{Z}_{3}$, to which $\mathrm{M}_{2}$ Must Continuelculations Assuming Solution A (Crustal Model A) |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{z}_{2}=120 \mathrm{~km}$ |  |  | $\mathrm{z}_{2}=130 \mathrm{~km}$ |  |  | $\mathrm{z}_{2}=140 \mathrm{~km}$ |  |  |
|  | $)=8$. | $\frac{\mathrm{km}}{\mathrm{sec}}$ |  | $=8$. | $\frac{\mathrm{km}}{\mathrm{sec}}$ |  | $)=8$ | $\frac{\mathrm{km}}{\mathrm{sec}}$ |
| $\begin{gathered} \left.\mathrm{M}_{2}-1\right) \\ \left(\mathrm{sec}^{-1}\right) \\ \hline \end{gathered}$ | $\begin{gathered} Z_{3} \\ (\mathrm{~km}) \\ \hline \end{gathered}$ | $\begin{gathered} V_{p}\left(Z_{3}\right) \\ (\mathrm{km} / \mathrm{sec}) \end{gathered}$ | $\begin{gathered} \mathrm{M}_{2}-1 \\ \left(\mathrm{sec}^{-1}\right) \end{gathered}$ | $\begin{array}{r} Z_{3} \\ (\mathrm{~km}) \\ \hline \end{array}$ | $\begin{gathered} V_{p}\left(z_{3}\right) \\ (\mathrm{km} / \mathrm{sec}) \end{gathered}$ | $\begin{gathered} \mathrm{M}_{2} \\ \left(\mathrm{sec}^{-1}\right) \end{gathered}$ | $\begin{gathered} \mathrm{Z}_{3} \\ (\mathrm{~km}) \\ \hline \end{gathered}$ | $\begin{gathered} V_{p}\left(Z_{3}\right) \\ (\mathrm{km} / \mathrm{sec}) \end{gathered}$ |
| 0.016 | 147.4 | 8.67 | 0.019 | 152.6 | 8.68 | 0.05 | 144.5 | 8.49 |
| 0.02 | 130.4 | 8.44 | 0.03 | 137.2 | 8.47 | 0.1 | 141.9 | 8.46 |
| 0.03 | 124.8 | 8.38 | 0.05 | 133.4 | 8.42 | 0.2 | 140.9 | 8.45 |
| 0.1 | 121.0 | 8.34 | 0.1 | 131.6 | 8.42 | 1 | 140.2 | 8.44 |
| 0.2 | 120.5 | 8.33 | 0.2 | 130.7 | 8.38 | 2 | 140.1 | 8.43 |
| 1 | 120.1 | 8.32 | 0.3 | 130.4 | 8.38 | 3 | 140.1 | 8.43 |
| 2 | 120.0 | 8.32 | 1 | 130.1 | 8.38 | 4 | 140.0 | 8.43 |
| 3 | 120.0 | 8.32 | 2 | 130.1 | 8.38 | 10 | 140.0 | 8.43 |
| 4 | 120.0 | 8.32 | 3 | 130.0 | 8.37 |  |  |  |
| 10 | 120.0 | 8.32 | 10 | 130.0 | 8.37 |  |  |  |

(pənutquoอ) $x$ әтqе山


is purely a function of $M_{2}$ :

$$
V_{p}\left(Z_{3}\right)=V_{p}\left(Z_{2}\right)+M_{2} x\left[Z_{3}\left(M_{2}\right)-Z_{2}\right]
$$

The physical situation corresponding to the limit as $M_{2}$ tends to infinity is a velocity discontinuity, from which $P_{2}$ would be a total reflection. Accordingly, from Table $X_{\text {, }}$ $\lim _{M_{2} \rightarrow \infty} Z_{3}\left(M_{2}\right)=Z_{2}$ and $\lim _{M 2 \rightarrow \infty} V_{p}\left(Z_{3}\right)$ exists and is greater than $V_{p}\left(Z_{2}\right)$. For example, for solution $A$ and $Z_{2}=130$ kilometers, $V_{p}\left(Z_{2}\right)=8.25 \mathrm{~km} / \mathrm{sec}$ and $\lim _{M 2 \rightarrow \infty} V_{p}\left(Z_{3}\right)=8.37 \mathrm{~km} / \mathrm{sec}$. For the entire range of solutions, the difference between $\lim _{M_{2} \rightarrow \infty} V_{p}\left(Z_{3}\right)$ and $\mathrm{V}_{\mathrm{p}}\left(\mathrm{Z}_{2}\right)$ is as small as $0.09 \mathrm{~km} / \mathrm{sec}$ and as large as 0.26 $\mathrm{km} / \mathrm{sec}$. When $\mathrm{Z}_{3}-\mathrm{Z}_{2}$ is large, then the difference $V_{p}\left(Z_{3}\right)-V_{p}\left(Z_{2}\right)$ is also large. For the entire range of solutions, the $P_{2}$ event can be explained by a difference between $Z_{3}$ and $Z_{2}$ as large as about 25 kilometers and a corresponding difference between $\mathrm{V}_{\mathrm{p}}\left(\mathrm{Z}_{3}\right)$ and $\mathrm{V}_{\mathrm{p}}\left(\mathrm{Z}_{2}\right)$ as large as about $0.45 \mathrm{~km} / \mathrm{sec}$.

Uncertainty of crustal structure was seen to have a pronounced effect. on the determination of the $P$ wave velocity distribution in the upper mantle based on $P_{1}$ arrival times. However, determination of deeper velocity structure, based on the $P_{2}$ event, is only slightly affected by this uncertainty.

Observed arrival times of $\mathrm{P}_{2}$ and theoretical $\mathrm{P}_{2}$ arrival
times for solution $A, Z_{2}=130$ kilometers, and $M_{2}=0.03$ $\sec ^{-1}$ are listed in Table XI.

## Table XI

## Arrival Times of $\mathrm{P}_{2}$

(Theoretical arrival times are given for Solution A, $\left.Z_{2}=130 \mathrm{~km}, M_{2}=0.03 \mathrm{sec}^{-1}\right)$

| Station | $\underset{(\mathrm{km})}{\text { Distance }}$ | Observed Arrival <br> Time (sec) | Theoretical Arrival Time (sec) |
| :---: | :---: | :---: | :---: |
| S2 | 830.6 | 110.56 | 111.26 |
| S3 | 898.4 | 119.11 | 119.14 |
| S4 | 979.9 | 128.73 | 128.68 |
| S5 | 1030.4 | 134.70 | 134.62 |
| S6 | 1080.6 | 140.61 | 140.54 |
| S7 | 1229.4 | 158.89 | 158.14 |
| S8 | 1284.3 | 165.06 | 164.65 |

## Incomplete Triplication

$P_{1}$ and $P_{2}$ have been considered to be part of a triplication (segments $W X$ and $X Y$ given in Fig. 8. There is, however, no event, $P_{3}$ say, observed on the records which corresponds to segment $Y Z$ in Fig. 8. The absence of a $\mathrm{P}_{3}$ event could be the result of the existence of a low velocity zone below $Z_{3}$ or a zone of very low gradient below $Z_{3}$. There is no direct evidence for a low velocity layer, but it is feasible that a zone of low gradient could produce a $P_{3}$ event of very small amplitude such that it would not be observed. For example, for solution $A$ and $Z_{2}=130$ kilometers, $M_{2}=2.0$ seconds ${ }^{-1}$, a velocity gradient of $1.0 \times 10^{-6}$ seconds ${ }^{-1}$, existing below $Z_{3}$, would produce a $P_{3}$ event such that the amplitude ratio $P_{1} / P_{3}$ is approximately 4 for all distances of the experiment. The ratio of $\mathrm{P}_{1}$ amplitude to noise amplitude, as stated in Chapter III, was found to be about 2.7. Thus, the $P_{3}$ event would be buried in noise.

Figure 10 is a reduced time-distance plot showing observed values and the theoretical graph for solution $A$, $Z_{2}=130$ kilometers, and $M_{2}=0.03$ seconds ${ }^{-1}$.

## Other Arrivals

Arrivals, other than $\mathrm{P}_{2}$, which occur within about 5 seconds after $P_{1}$ may be part of minor triplications (Green and Steinhart, 1962). Since the records are of

varying quality and the station spacing is large (up to 150 kilometers), it is difficult to trace events, which may be part of minor triplications, from record to record.

The $\overline{\mathbf{P}}$ Phase
The $\overline{\mathbf{P}}$ phase is present on all of the records; it is very strong on all records except S 8 . Figures 12 a and 12 b show $\bar{P}$ on records $S 2$ and $S 8$ respectively. Velocities and arrival times for the $\overline{\mathrm{P}}$ phase are given in Table XII. The velocity given is the distance divided by the arrival time.

The $\bar{P}$ phase arrives at times expected for the direct wave, $P_{g}$. However, for distances of this experiment, $P_{g}$ theoretically should not be visible. The presence of $\bar{P}$ indicates a velocity gradient in the crust.

Results from Project Early Rise

Green and Hales (1968) have interpreted records from Project Early Rise to determine upper mantle structure in the Central United States. Two Early Rise models are proposed. For Model 1 , velocity increases slowly below the Moho ( 50 km depth); a rapid increase in velocity gradient occurs at 89 km (the velocity increases by $0.26 \mathrm{~km} / \mathrm{sec}$ ); below 89 km , the velocity gradient is low. This model is similar in form to the models $A, B$, and $C$. Model 2 is similar to Model 1 down to a depth of about 134 km ; at this depth Model 2 includes a low velocity layer 25 km thick. However, observations explained by the low velocity layer may also be explained by lateral velocity variation.

# Table XII 

$\overline{\mathrm{P}}$ Phase

| Station | Distance <br> $(\mathrm{km})$ | Arrival Time <br> $(\mathrm{sec})$ | Velocity <br> $(\mathrm{km} / \mathrm{sec})$ |
| :--- | :---: | :---: | :---: |
| S1 | 793.2 | 131.75 | 6.02 |
| S2 | 830.6 | 138.38 | 6.00 |
| S3 | 898.4 | 149.97 | 5.99 |
| S4 | 979.9 | 164.63 | 5.95 |
| S5 | 1030.4 | 174.03 | 5.92 |
| S6 | 1080.6 | 180.06 | 6.00 |
| S7 | 1229.4 | 207.47 | 5.93 |
| S8 | 1284.3 | 214.65 | 5.98 |










$$
1 \Omega-\frac{n}{N}
$$

$$
\underset{\sim}{\sim} \underset{\sim}{\sim}
$$

H 犬 N N
O

$$
\text { Fig. 12b. } \overline{\mathrm{P}} \text { Phase, Record S8. Distance }=1284.3 \mathrm{~km} .
$$

## CHAPTER V

## CONCLUSIONS

Digital filtering techniques offered no improvement over simple analog playbacks of the seismic records obtained; the VLF-2 system, which is designed for long range refraction experiments, was found to be effective in increasing the signal to noise ratio of the seismic data.
$P_{1}$, the first arrival, arrives at times in accordance with a velocity function that increases linearly and slowly with depth below the base of the crust. Uncertainty of crustal structure, however, produces uncertainty in the velocity at the base of crust, $\mathrm{V}_{\mathrm{m}}$, and the velocity gradient within the upper mantle. An average crustal structure for interior plains and plateaux suggests a value of $8.10 \pm 0.5$ $\mathrm{km} / \mathrm{sec}$ for $\mathrm{V}_{\mathrm{m}}$, and a gradient between $0 \mathrm{sec}^{-1}$ and about $0.003 \mathrm{sec}^{-1}$. Arrival times of a second event, $\mathrm{P}_{2}$, and observed $P_{2} / P_{1}$ amplitude ratios suggest a rapid increase in velocity gradient occurring between depths of about 120 kilometers and 150 kilometers. The incomplete triplication formed by $P_{1}$ and $P_{2}$ suggests the existence of a zone of low velocity gradient below the rapid increase. Thus, it is not necessary to explain the incomplete triplication by the existence of a low velocity zone.

## APPENDIX I

## RAY THEORY AND POSITIVE LINEAR VELOCITY-DEPTH GRADIENTS

## Travel-time and Distance Equations

It was found that for the distance range of this study the flat earth approximation was inaccurate. The travel time and distance of a ray which travels between radii $r_{1}$ and $r_{2}$ in a spherically stratified earth are given by Bullen (1963).

$$
\begin{array}{cc}
\Delta=p \int_{r_{2}}^{r_{1}} r^{-1}\left(n^{2}-p^{2}\right)^{-\frac{1}{2}} d r & \ldots I \cdot 1 \\
\Delta=\int_{r_{2}}^{r_{1}} n^{2} r^{-1}\left(n^{2}-p^{2}\right)^{-\frac{1}{2}} d r & \ldots I \cdot 2 \\
p=\frac{r}{V} \sin \alpha & \ldots I \cdot 3
\end{array}
$$

where,
$\Delta=$ angular distance travelled by ray
$T=$ travel time of ray
$r=$ distance from center of earth to point on ray path
$\alpha=$ angle between direction of ray path and radius vector
$v=$ velocity (a function of $r$ only)
$p=$ ray parameter (constant for each ray)
$n=r / V$

An earth model consisting of any number of spherical shells, for which the velocity in the $i$ th shell is given by equation I. 4 , has been considered.

$$
V=m_{i} r+b_{i} \quad \ldots I \cdot 4
$$

where $m_{i}$ and $b_{i}$ are constants. For positive velocity-depth gradients, $m_{i} \leq 0$ and therefore $b_{i} \geq 0$. The following definitions are useful;

$$
\begin{aligned}
r_{i}= & \text { radial coordinate to the top of the } i^{\text {th }} \text { shell } \\
v_{i}= & \text { velocity at the top of the } i \text { th shell } \\
\Delta_{i}= & \text { angular distance travelled by ray through the } \\
& i \text { th shell } \\
T_{i}= & \text { travel time of ray through the } i^{\text {th }} \text { shell }
\end{aligned}
$$

From equations $I \cdot 1, I \cdot 2$, and $I \cdot 4, \Delta$ and $T$ for the $i^{\text {th }}$ shell are,

$$
\begin{gathered}
\Delta_{i}=2 p \int_{r_{i+1}}^{r_{i}} r^{-1}\left[\frac{r^{2}}{\left(m_{i} r+b_{i}\right)^{2}}-p^{2}\right]^{-\frac{1}{2}} d r \quad \ldots I \cdot 5 \\
T_{i}=2 \int_{r_{i+1}}^{r_{i}} \frac{r}{\left(m_{i} r+b_{i}\right)^{2}}\left[\frac{r^{2}}{\left(m_{i} r+b_{i}\right)^{2}}-p^{2}\right]^{-\frac{1}{2}} d r \quad \ldots I \cdot 6
\end{gathered}
$$

Stewart (1968) solves equations I•5 and I•6. The following solutions are based on those given by Stewart, but differ in the following way; expressions of the form $\ln (x)$ have been changed to $\ln (|x|)$. For $m<0$, the solutions of $I \cdot 5$ and $I \cdot 6$ depend upon the value of $C=1-m_{i}{ }^{2} p^{2}$.
$C<0$

$$
\begin{aligned}
& \frac{\Delta_{i}}{2}=\left\{\frac{m_{i} p}{\sqrt{-c}} \arcsin \left[\frac{-C r}{p b_{i}}+m_{i} p\right]+\arcsin \left[-m_{i} p-\frac{p b_{i}}{r}\right]\right\} r_{i} \\
& r_{i+1} \\
& \frac{T_{i}}{2}=\left\{\frac{-1}{m_{i}}\left[\frac{1}{\sqrt{-C}} \arcsin \left[\frac{-C V+b_{i}}{-p m_{i} b_{i}}\right]+\ln \left(\left|\frac{-m_{i} \sqrt{r^{2}-p^{2} V^{2}}+b_{i}}{V}-1\right|\right)\right]\right\} r_{i}, v_{i} \\
& r_{i+1}, v_{i+1} \\
& C=0
\end{aligned}
$$

$$
\left.\begin{array}{c}
\frac{\Delta_{i}}{2}=\left\{-\sqrt{-1-2 m_{i} r}\right. \\
b_{i}
\end{array} \arcsin \left[1+\frac{b_{i}}{m_{i} r}\right]\right\}_{i}^{r_{i}} r_{i+1} .\left(\sqrt{b_{i}} \frac{T_{i}}{2}=\left\{\frac{1}{m_{i}}\left(\sqrt{1-\frac{2 V}{b_{i}}}+\ln \left(\left|\left(\sqrt{1-\frac{2 V}{b_{i}}}-1\right) /(\sqrt{1-2 V}+1)\right|\right)\right)\right\}_{V_{i+1}}^{V_{i}} .\right.
$$

$$
c>0
$$

$$
\begin{array}{r}
\frac{\Delta_{i}}{2}=\left\{\frac{m_{i} p}{\sqrt{c}} \ln \left(\left|\sqrt{r^{2}-p^{2} V^{2}}+r \sqrt{c}-\frac{m_{i} b_{i} p^{2}}{\sqrt{c}}\right|\right)+\arcsin \left[-m_{i} p-\frac{p b_{i}}{r}\right]\right\}{ }_{r_{i}}, v_{i} \\
r_{i+1} \\
v_{i+1}
\end{array}
$$

$$
\frac{T_{i}}{2}=\left\{\frac{-1}{m_{i} \sqrt{C}} \ln \left(\left|-m_{i} \sqrt{r^{2}-p^{2} V^{2}}+V \sqrt{C}-\frac{b_{i}}{\sqrt{C}}\right|\right)-\frac{1}{m_{i}} \ln \left(\left|\frac{-m_{i} \sqrt{r^{2}-p^{2} V^{2}}+b_{i}}{V} 1\right|\right)\right\} r_{i,} r_{i}
$$

$$
v_{i+1}
$$

When $m_{i}=0, \Delta_{i}$ and $T_{i}$ are,

$$
\left.\frac{\Delta_{i}}{2}=\left\{\operatorname{arcos}\left[p b_{i} / r\right]\right\}\right\}_{i}^{r_{i+1}} \quad \frac{T_{i}}{2}=\left\{\frac{1}{b_{i}} \sqrt{r^{2}-p^{2} b_{i}^{2}}\right\}^{r_{i}} r_{i+1}
$$

For a ray which bottoms in the $i^{\text {th }}$ layer, $\Delta_{i}$ and $T_{i}$ are found from one of the above sets of equations by substituting $r_{B}$ and $V_{B}$ for $r_{i+1}$ and $V_{i+1}$.

$$
\begin{aligned}
r_{B}= & \text { the radial coordinate at the deepest point of } \\
& \text { penetration. } \\
V_{B}= & \text { velocity at } r_{B} .
\end{aligned}
$$

The total travel time and distance for a ray which bottoms in the $n^{\text {th }}$ spherical shell are:

$$
\Delta=\sum_{i=1}^{n} \Delta_{i} \quad T=\sum_{i=1}^{n} T_{i}
$$

## Amplitude Ratios (Geometric Spreading)

For an energy source at the earth's surface, the effect of geometric spreading on vertical amplitude is given by (Bullen, 1963),
$A^{2} \alpha \frac{I \tan ^{2} e \sec ^{2} e\left(1+3 \tan ^{2} e\right)^{2}}{n^{2} \sin \Delta\left(\tan ^{2} e-\sin ^{2} e\right)^{\frac{1}{2}}\left\{4 \tan e \tan f+\left(1+3 \tan ^{2} e\right)^{2}\right\}^{2}}\left|\frac{d^{2} T}{d \Delta^{2}}\right| \quad I \cdot 7$
where:

$$
\begin{aligned}
& \mathrm{A}=\text { vertical amplitude at recording site } \\
& I=\text { power/unit solid angle at source } \\
& e=\text { angle of emergence at recording site } \\
& \mathrm{n}=r / V \text { at surface } \\
& \cos ^{2} f=\cos ^{2} e / 3
\end{aligned}
$$

Since $d^{2} T / d \Delta^{2}=1 / d \Delta / d p, A$ can be calculated for various
rays arriving at the same distance; thus, vertical amplitude ratios can be determined.

## APPENDIX II

## PROGRAM "RAY" DESCRIPTION

1. Identification

Title: Calculations of time and distance for rays which travel in an earth model consisting of any number of spherical shells, in each of which velocity increases linearly with depth.

Programmer: Allan Bates
Date: September, 1970
Language: Fortran IV
2. Purpose

To show the effects of spherical shells, in which velocity increases linearly with depth, on the timedistance relation.
3. Usage

Operational Procedure: The main program reads the
input data. Subroutine "Ray" calculates time-distance
tables.
Input Parameters:
$N M=$ number of models for which tables are to be calculated.

NN $=$ number of layers +1 for the models.
$(Z(I), I=1, N N)=$ depth to the top of the $i^{\text {th }}$ layer (kilometers). $Z(1)$ must always be zero. (VC(I), I $=1, N N-1)=$ velocity at the top of the $i^{\text {th }}$ layer (km/sec).
$(M(I), I=1, N N-1)=$ linear velocity gradient in the $i^{\text {th }}$ layer (seconds ${ }^{-1}$ ). $M(I)$ must always be less than or equal to zero.

## Calculated Parameters

( $\mathrm{RC}(\mathrm{I}), \mathrm{I}=1, \mathrm{NN}-1)=$ distance from centre of earth to top of $i^{\text {th }}$ layer (kilometers).
(RD(I), I $=1, \mathrm{NN}-1)=$ distance from center of earth to bottom of $i^{\text {th }}$ layer (kilometers).
$(\operatorname{VCX}(I), I=I, N N-I)=$ velocity at the bottom of the $i^{\text {th }}$ layer ( $\mathrm{km} / \mathrm{sec}$ ).

For each ray, the following values are calculated, DB $=$ depth at which ray bottoms (kilometers). $\mathrm{RB}=$ distance from center of earth to point at which ray bottoms (kilometers). $\mathrm{VB}=$ velocity at DB (km/sec). IB $=$ number of layer in which ray bottoms. (DEL(I), $I=1, I B)=$ half the distance (degrees) travelled in the $i^{\text {th }}$ layer.
$(T I), I=1, I B)=$ half the travel time in the $i^{\text {th }}$ layer (seconds).

DIST $=$ total distance travelled (kilometers). TIME $=$ total travel time (seconds).

```
AVEL = apparent velocity of time-distance relation at the distance at which the ray emerges ( \(\mathrm{km} / \mathrm{sec}\) ).
```


## 4. Comments

The program is used for spherical shells for which the velocity, $V$, in the $i^{\text {th }}$ shell is,

$$
\mathrm{V}=\mathrm{M}(\mathrm{I}) \times \mathrm{R}+\mathrm{B}(\mathrm{I}) \quad \ldots . . \mathrm{II} .1
$$

where,
$R=$ distance from center of earth
$B(I)=$ constant $(\mathrm{km} / \mathrm{sec})$
When velocity is expressed as a function of depth, equation (II.1) becomes

$$
V=V C(I)-M(I) \times(Z-Z(I)) \quad \ldots I I .2
$$

where $\mathrm{Z}=$ depth (kilometers).
Velocity must increase with depth and thus $M(I) \leq 0$. Equation (II.2) is convenient for determining input parameters. Within the program, equation (I.1) is used for calculations.

Following is an example of "Ray". Calculated values of DB, DIST, and TIME are given in the output. The object program required 36 k bytes of storage space. The central processing unit time for calculations involving over 600 rays was 0.22 min .; 7.34 seconds of this time was used for the compile step.

```
    OOUBLE PRECISION Z(10),VC(9),RC(9),RD(9),VCX(9),B(9),M(9),BN(9),
    1DEL(9),T(9)
                                    72
    PROGRAM RAY GOMPUTES TIME-DISTANCE TABLES FOR EARTH
    MODELS CONSISTING OF ANY NUMBER OF SPHERICAL SHELLS.
    VELJGITY INCREASES LINEARLY UITH DEPTH
                INPUT AS FOLLOWS,
                NM=NUMBER OF MODELS
                NN = NUMBER OF LAYERS+1
                (Z(I),I=1,NN)=DEPTH TO TOP OF ITH LAYER
                (Z(1) MUST ALWAYS BE ZERO)
                (VC(I),I=1,NN-1)=VELOCITY AT TOP OF ITH LAYER
                (M(I),I=1,NN-1)=LINEAR GRADIENT IN ITH LAYER
                    (M(I) MUST BE LESS THAN OR EQUAL TO ZERO)
            VELOCITY IN EACH LAYER IS,
                    V=VC(I)-M(I)*(Z-Z(I)),WHERE,V=VELOCITY,Z=DEPTH
            OUTPUT,
                A TABLE RELATING DEPTH OF PENETRATION,
                DISTANCE TRAVELLED,AND TRAVEL TIME
            THE MAIN PRDGRAM READS THE INPUT DATA .CALCULATIONS
            ARE DONE BY SUBRDUTINE RAY
    101 FORMAT (2I5)
    102 FORMAT(5F10.2)
    103 FORMAT(5F10.5)
        READ(5,101) NM,NN
        LL=0
        J=NN-1
        REAO(5,102) (Z(I), I=1,NN)
        READ(5,102) (VC(I),I=1,J)
        READ(5,103) (M(I),I=1,J)
490 CALL RAYINN,J,Z,VC,RC,RD,VCX,B,M,BN,DEL,T)
    LL=LL+1
    IF(LL.LT.NM) GG TO 490
9 6 ~ C O N T I N U E ~
    CALL EXIT
    END
```

SUBROUTINE RAY(NN,J,Z,VC,RC,RD,VCX,B,M,BN,DEL,T)
DOUBLE PRECISION Z(NN), VC(J), RC(NN), RD(J),VCX(J),B(J),M(J),
73
1BN(J),RB, $V B, P, D B, D E L(J), T(J), A, F, P H I, D I S T, T I M E, P I E, A V E L, C$
DEFINITION OF PARAMETERS,
(RC(I), $I=1, N N-1)=D I S T A N C E$ FROM CENTER OF EARTH TO TOP OF ITH LAYER
(RD(1), $I=1, N N-1)=D I S T A N G E$ FROM CENTER OF EARTH TO BOTTOM OF ITH LAYER
$(V C X(I), I=1, N N-1)=V E L O C I T Y$ AT THE BOTTOM OF THE
ITH LAYER
DB=DEPTH AT WHICH RAY BOTTOMS
$R B=D I S T A N C E$ FROM CENTER OF EARTH TO POINT AT
WHICH RAY BOTTOMS
$V B=V E L O C I T Y$ AT DB
AVEL = APPARENT VELOCITY OF RAY WHICH BOTTOMS AT DB (DEL $(1), I=1, N N-1)=$ HALF CONTRIBUTION OF ITH LAYER TO DISTANCE. (DEL(I) IS GALCULATED IN DEGREES)
$(T(I), I=1, N N-1)=H A L F$ CONTRIBUTION OF ITH LAYER TO TRAVEL TIME
$P=$ CONVENTIONAL RAY PARAMETER (BULLEN, 1963)
THE SUBROUTINE USES THE FOLLOWING RELATION FOR VELOCITY
IN THE ITH LAYER,
$V=M(I) * R+B(I)$
WHERE R=DISTANCE FROM CENTER OF EARTH
200 FORMAT ( $1 \mathrm{H}, 15 \mathrm{X}$, NUMBER OF LAYERS $=$, , I 3)

202 FORMAT(14, /, 15X,F8. $2,6 X, F 13.2,14 X, F 11.5)$
203 FORMAT(IH, 15 X, 'PENETRATION(KM)', 8 X, "DISTANCE(KM)', $10 X$, TIME(SEC) 1')
204 FORMAT (1H, 10X,F15.2,10X,F12.2.10X,F9.2)
205 FORMAT (1H, ////)
WRITE INPUT DATA
WRITE $(6,200) \mathrm{J}$
WRITE $(6,201)$
DO $50 \quad \mathrm{I}=1, \mathrm{~J}$
WRITE (6,202) Z(I),VC(I),M(I)
$5)$ CONTINUE
WRITE $(6,205)$
WRITE $(6,203)$
GALCULATE RADIUS OF EARTH AT 50.75 DEGREES LATITUDE RADIUS OF EARTH=RC(1)
$\mathrm{PIE}=3.141592653589793200 / 2.000$
$\mathrm{A}=6378.38800$
$F=1 . D 0 / 297 . D 0$
$\mathrm{PHI}=(50.7500 * 3.141592653589793200) /(180.000)$
$\operatorname{RC}(1)=A *(1 . D O-F *(D S I N(P H I) * D S I N(P H I))$
$1+(5.00 / 8.00) * F * F \div(D S I N(2 . D O * P H I)) *(D S I N(2 . D O * P H I))$
CALCULATE RC(I), RDII), VCX(I), B(I)
DO $20 \quad I=1, J$
$R C(I+1)=R C(1)-2(I+1)$
$B(I)=V C(I-M(I) * R C(I)$
$V C X(I)=V C(I)-M(I) *(Z(I+1)-Z(I))$
$B N(I)=B(I) / D A B S(B(I))$
$R O(I)=R C(1)-2(1+1)$
20
GONTINUE
INITIATE VALUE OF RAY PARAMETER
$P=R C(2) / V C(2)+0.049 D 0$
DO $1 \mathrm{I}=1, \mathrm{~J}$
$I B=I$
$I B 2=I+1$
$R B=(P * B(I)) /(1 . D 0-P * M(I))$
$V B=M(I) \div R B+B(I)$
IF(RB.GE.RC(I)) GO TO 70
IF(RB.LE.RC(I).AND.RB.GE.RC(IB2)) GOTO 2
GO TO 1
$R B=R C(I)$
$V B=V C$ (I)
IREF=1
$I B=I B-1$
WRITE $(6,71)$ IB
71 FORMAT (1H, 'REF FROM', 13)
GO TO 2
1 CONTIMUE
$2 \quad D B=R C(1)-R B$
$A V E L=V B * R C(1) / R B$
IF RAY BOTTOMS BELOW REGION OF INTEREST, END CALCULATION
FOR PARTICULAR MODEL
IF (RB.LT.RC(NN)) GO TO 300
DIST $=0.000$
TIME $=0.000$
FInal Value of dist is total distance travelled
FINAL VALUE OF TIME IS TOTAL TRAVEL TIME
DO LOOP WHICH GALCULATES DELII) AND T(I)
DO 4 I $=1$, IB
$C=1.0 D O-M(I) * M(I) * P * P$
IFIIREF.EQ.1) GO TO 60
IF(I.NE.IB) GO TO 60
SPECIAL EQUATIONS FOR LAYER IN WHICH RAY BOTTOMS
IF(M(I).EQ.O.ODO) GO TO 45
IF(C) $46,47,48$
45
$D E L(I)=(D A R S I N(-C * R C(I) /(P * D A B S(B(I)))+M(I) * P * B N(I))\} * M(I) *$
$1 P /((-C) * * .5 D 0)$
$1+(\operatorname{DARSIN}(-M(I) * P * B N(I)-P * D A B S(B(I)) / R C(I))) * B N(I)$
$1+(M(1) * P /((-C) * 5 D O)+1 . D 0) * P I E$
$T(I)=((D A R S I N(-C * V C(I)+B(I)) /(P * D A B S(M(I) * B(I))))$
$1 /((-C) * * 5 D)+(D L O G(D A B S((D A B S(M(I)) *(\operatorname{RC}(I) * R C(I)-P * P * V C I I) * V C(I$
1)) $* * 5001$
$1+\operatorname{DABS}(B(I)) /(\operatorname{VC}(I))-B N(I))) * B N(I)) / D A B S(M(I))$
$1-(P I E /((-5) * * 500)+D L O G(D A B S(-M(I) * R B / V B))) / D A B S(M(I))$
GO TO 9

$1+(\operatorname{DARS} \operatorname{si}(-I M(I) * B(I)) /(D A B S(M(I) * B(I)))$
$1-D A B S(B(I) / M(I)) / R C(I))) * B N(I)+P I E$
$T(I)=-((1 . D 0-2 . D 0 * V C(I) / B(1)) * * * 500$
$1+\operatorname{LOG}(D A B S((1 . \operatorname{DO}-2 . D O * V C(I) / B(I)) * * 5 D 0-1 . D 0) /$
$1(1.00-2.00 * V G(I) / B(I)) * 500+1 . D 0) 1) *(B N(I) / D A B S(M(I))$
60 109
$48 \mathrm{DEL}(\mathrm{I})=(\mathrm{OLOG}(\mathrm{DABS}($ (RC(I)*RC(I)-P*P*VC(I)*VC(I))**.5D0


```
    1(こ**.5DO))+(DARSIN(-M(I)*p*BN(I)-P*DABS(B(I))/RC(I)))
    1*BN(I)-(DLOG(DABS (P*B(I)/(C**.5D0))))*M(I)*P/(C**.5DO)+PIE
        T(I)=(DLDG(DABS(DABSIM(I))*((RC(I)*RC(I)-P*P*
        IVC(I)*VC(I))**.5D0)+VC(I)*(C**.5DO)-B(1)/(C**.500))))
        1/(DABS(M(I))*(C**.5DO))
        1+(DLJG(DABS((DABSIM(I))*((RC(I)*RC(I)-P*P*VC(I)*VC(I))**.5DO)+
        1DABS(B(I)))/(VC(I))-BN(I))) %(BN(I)/(DABS(M(I))))
        1-(DLOG(DABS(M(I)*P*B(I)/(C**5DO))))/(DABS(M(T))*(C**.5DO))
        1-(DLOG(DABS(-M(L)*P)))*BN(I)/(DABS(M(I)))
        GO TO 9
    45 DEL(I)=DARCOS(P*B(I)/RC(I))
        T(I)=((RC(I)*RC(I)-P*P*B(I)*B(I))**.5DO)/B(I)
        GO TO 9
C
    60 IF(M(I).EQ.O.ODO) GO TO 5
        IF(C) 6,7,8
    5 DEL(I)=(DARSIN(-C*RC(I)/(P*DABS(B(I)))+M(I)*P*BN(I))-
        IDARSIN(-C*RD(I )/(P*DABS(B(I)))+M(I)*P*BN(I)))*M(I)*P/((-G)**.5DO)
        1+(DARSIN(-M(I)*P*BN(I)-P*DABS(B(I))/RC(I))
        1-DARSIN(-M(I)*P*BN(I)-P*DABS(B(I))/RD(I )))*BN(I)
        T(I)=((DARSIN((-C*VC(I)+B(I))/(P*DABS(M(I)*B(I)))))/((-C)**.5DO)
        1+(DLOG(DABS((DABS(M(I))*((RC(I)*RC(I)-P*P*VC(I)*VC(I))**.5D0)
        1+DABS(B(I)))/(VC(I))-BN(I))))*BN(I))/(DABS(M(N)))
        1-((OARSIV({-C*VCX(I)+B(I))/(P*DABS(M(I)*B(I)))))/((-C)**.5DO)
        1+(DLDG(DABS([DABS (M(I))*((RD(I)*RD(I)-P*P*VCXII)*VCX(I))**5DO)
        I+DABS(B(I)))/(VCX(I))-BN(I))))*BN(I))/(DABS(M(I)))
        GO TO 9
    7 DEL(I)=-((-1.DO-2.DO*M(I)*RC(I)/B(I))**.5DO)*BN(I)
        1+(DARSIN(-1M(I)*B(I))/(DABS(M(I)*B(I)))
    1-DABS(B(I)/M(I))/RC(I)))*BN(I)
    1+({-1.DO-2.DO*M(I)*RD(I)/B(I))**.5DO)*BN(I)
    l-(DARSIVI-(M(I)*B(I))/(DABS(M(I)*B(I)))-DABS(B(I)/M(I))
    1/RD(I )|)*BN(I)
        T(I)=-((1.DO-2.DO*VC(I)/B(I))**.5DO+DLOG(DABS((1).DO-2.DO*VG(I)/
    1B(I))**.5DO-1.DO1/
    1((1.DO-2.DO*VC(I)/B(I))**.5DO+1.DO))))*(BN(I)/DABS(M(I)))
    1+((1.DO-2.OD*VCX(I)/B(I))**.5DO+
    1DLOG(DABS(1(1.DO-2.DO*VCX(1)/B(I))**.5D0-1.DO)
    1/((1.DO-2.DO*VCX(I)/B(I))**.5DO+1.DO))))*(BN(I)/DABS(M(I)))
        GO TO 9
3 DELII)=(DLOG(DABS(IRCII)*RC(I)-P*P*VC(I)*VC(I))**.5D0
    1+RC(I)*(C**.5DO)-M(I)*B(I)*P*P/(C**.5DO)))
    1-DLOG(DABSI(RD(I)*RD(I)-P*P*VCX(I)*VCX(I)]**5DO
    1+RD(I)*(C**.5DO)-M(I)*B(I)*P*P/(C**.5DO))))*(M(I)*P/(G**.5D0))
    1+(DARSIV(-M(I)*P*EN(I)-P*DABS(B{I))/RC(I))
    1-DARSIN(-M(I)*P*BN(I)-P*DABS(B(I))/RD(I )))*BN{I)
        T(I)={DLOG(DABS (DABS{M(I))*({RC(I)*RC(I)-P*P*VC(I)*VC(I))***5D0)+
    IVC(I)*(5**.5D0)-B(I)/(C***5DO)))
    1-DLOG(DABS (DABS(M(I))*(IRD(I)*RD(I)-P*P*VCX(I)*VCX(I))***500)+
    IVCX(I)*(5**.5DO)-B(I)/(C**.5D0))11/(DABS(M(I))*(C**.5D0))
    1+(DLJG(DABS(IDABS(M(I))*((RC(I)*RC(I)-P*P*VC(I)*VC(I))**.5DO)
    1+DABS(B(1)) //(VC(I))-BN(I)))
    1-DLOG(DABS(IDABS(M(I))*((RD(I)*RD(I)-P*P*VCX(I)*VCX(I))**.5DO)
    1+OABS(B(I))/(VGX(I))-BN(I))))*(BN(I)/(DABS(M(I))))
        GO TO 9
5 DEL(I)=DARCOS(P*B(I)/RC(I))-DARCOS(P*B(I)/RDII )
    T(I)=((RC(I)*RC(I)-P*P*B(I)*B(I))**.500
    1-(RD(I )*RD(I -P*P*B(I)*B(I))**500)/B(I)
```

9 DIST $=$ DIST+RC(1)*DEL (I) *2.DO
TIME=TIME+T(I)*2.DO
IF(I.NE.IB) GO TO 500
C BY RAY AND TRAVEL TIME OF RAY
WRITE 16,204 DB,DIST, TIME
500 CONTINUE
4 CONTINUE
GO TO 10
300 CONTINUE
RETURN
END

NUMBER OF LAYERS = 2
Z(I) $\quad$ VCM) (I) (KM/SEC)
$M(I)(1 / S E C)$
6.34
41.06
8.10
0.0
8.10
0.0
$-0.00170$

PENETRATION(KM)
41.06
41.24
41.41
41.59
41.76
41.93
42.11
42.28
42.46
42.63
42.80
42.98
43.15
43.33
43.50
43.68
43.85
44.02
44.20
44.37
44.55
44.72
44.89
45.07
45.24
45.42
45.59
45.77
45.94
46.12
46.29
46.46
46.64
46.81
46.99
47.16
47.34
47.51
47.69
47.86
48.03
48.21
48.38
48.56
48.73
48.91
49.08
49.26
49.43
49.61
49.78
49.95
rn $\quad$ n

| 61.86 | 778.56 | 103.39 |  |
| :---: | :---: | :---: | :---: |
| 62.04 | 781.40 | 103.74 |  |
| 62.21 | 784.24 | 104.08 |  |
| 62.39 | 737.06 | 104.42 | 78 |
| 62.56 | 789.87 | 104.77 |  |
| 62.74 | 792.67 | 105.11 |  |
| 62.92 | 795.46 | 105.45 |  |
| 63.09 | 798.23 | 105.78 |  |
| 63.27 | 801.00 | 106.12 |  |
| 63.44 | 803.76 | 106.46 |  |
| 63.62 | 806.50 | 106.79 |  |
| 63.79 | 809.24 | 107.12 |  |
| 63.97 | 811.96 | 107.45 |  |
| 64.14 | 814.68 | 107.78 |  |
| 64.32 | 817.38 | 108.11 |  |
| 64.50 | 820.08 | 108.44 |  |
| 64.67 | 822.76 | 108.77 |  |
| 64.85 | 825.44 | 109.09 |  |
| 65.02 | 828.10 | 109.42 |  |
| 65.20 | 830.75 | 109.74 |  |
| 65.38 | 833.41 | 110.06 |  |
| 65.55 | 836.04 | 110.38 |  |
| 65.73 | 838.67 | 110.70 |  |
| 65.90 | 841.29 | 111.02 |  |
| 66.08 | 843.90 | 111.34 |  |
| 66.25 | 846.50 | 111.65 |  |
| 66.43 | 849.09 | 111.97 |  |
| 66.61 | 851.68 | 112.28 |  |
| 66.78 | 854.25 | 112.60 |  |
| 66.96 | 856.82 | 112.91 |  |
| 67.13 | 859.37 | 113.22 |  |
| 67.31 | 861.92 | 113.53 |  |
| 67.49 | 864.46 | 113.84 |  |
| 67.66 | 867.00 | 114.14 |  |
| 67.84 | 869.52 | 114.45 |  |
| 68.01 | 872.04 | 114.76 |  |
| 68.19 | 874.55 | 115.06 |  |
| 68.37 | 877.05 | 115.36 |  |
| 68.54 | 879.54 | 115.67 |  |
| 68.72 | 882.03 | 115.97 |  |
| 68.89 | 884.50 | 116.27 |  |
| 69.07 | 886.97 | 116.57 |  |
| 69.25 | 889.44 | 116.87 |  |
| 69.42 | 891.89 | 117.17 |  |
| 69.60 | 894.34 | 117.46 |  |
| 69.77 | 896.78 | 117.76 |  |
| 69.95 | 899.21 | 118.06 |  |
| 70.13 | 901.64 | 118.35 |  |
| 70.30 | 904.06 | 118.64 |  |
| 70.48 | 906.47 | 118.94 |  |
| 70.65 | 908.87 | 119.23 |  |
| 70.83 | 911.27 | 119.52 |  |
| 71.01 | 913.66 | 119.81 |  |
| 71.18. | 916.04 | 120.10 |  |
| 71.36 | 918.72 | 120.39 |  |
| 71.54 | 920.79 | 120.67 |  |
| 71.71 | 923.16 | 120.96 |  |
| 71.89 | 925.51 | 121.25 |  |
| 72.06 | 927.86 | 121.53 |  |
| 72.24 | 930.21 | 121.82 |  |
| 72.42 | 932.55 | 122.10 |  |
| 72.59 | 934.88 | 122.38 |  |
| 72.77 | 937.20 | 122.66 |  |
| 72.95 | 939.52 | 122.94 |  |
| 73.12 | 941.83 | 123.23 |  |


| 30.30 | 353.35 | 15.91 |
| :---: | :---: | :---: |
| 50.48 | 557.58 | 76.43 |
| 50.65 | 561.78 | 76.94 |
| 50.83 | 565.94 | 77.45 |
| 51.00 | 570.06 | 77.96 |
| 51.18 | 574.14 | 78.46 |
| 51.35 | 578.20 | 78.95 |
| 51.53 | 582.21 | 79.44 |
| 51.70 | 586.20 | 79.93 |
| 51.88 | 590.15 | 80.41 |
| 52.05 | 594.07 | 80.89 |
| 52.23 | 597.96 | 81.37 |
| 52.40 | 601.82 | 81.84 |
| 52.58 | 605.65 | 82.30 |
| 52.75 | 609.45 | 82.77 |
| 52.93 | 613.22 | 83.23 |
| 53.10 | 616.97 | 83.69 |
| 53.28 | 620.69 | 84.14 |
| 53.45 | 624.38 | 84.59 |
| 53.63 | 628.04 | 85.04 |
| 53.80 | 631.69 | 85.48 |
| 53.98 | 635.30 | 85.93 |
| 54.15 | 638.90 | 86.36 |
| 54.33 | 642.47 | 86.80 |
| 54.50 | 646.01 | 87.23 |
| 54.68 | 649.54 | 87.66 |
| 54.85 | 653.04 | 88.09 |
| 55.03 | 656.52 | 88.52 |
| 55.20 | 659.97 | 88.94 |
| 55.38 | 663.41 | 89.36 |
| 55.55 | 666.83 | 89.77 |
| 55.73 | 670.22 | 90.19 |
| 55.90 | 673.60 | 90.60 |
| 56.08 | 676.96 | 91.01 |
| 56.25 | 680.29 | 91.42 |
| 56.43 | 683.61 | 91.82 |
| 56.60 | 686.91 | 92.22 |
| 56.78 | 690.19 | 92.62 |
| 56.95 | 693.46 | 93.02 |
| 57.13 | 696.70 | 93.42 |
| 57.30 | 699.93 | 93.81 |
| 57.48 | 703.14 | 94.20 |
| 57.65 | 706.34 | 94.59 |
| 57.83 | 709.51 | 94.98 |
| 58.00 | 712.67 | 95.36 |
| 58.18 | 715.82 | 95.75 |
| 58.35 | 718.95 | 96.13 |
| 58.53 | 722.06 | 96.51 |
| 58.70 | 72.16 | 96.89 |
| 58.88 | 728.25 | 97.26 |
| 59.05 | 731.32 | 97.64 |
| 59.23 | 734.37 | 98.01 |
| 59.41 | 737.41 | 98.38 |
| 59.58 | 740.43 | 98.75 |
| 59.76 | 743.45 | 99.11 |
| 59.93 | 746.44 | 99.48 |
| 60.11 | 749.43 | 99.84 |
| 60.28 | 752.40 | 100.20 |
| 60.46 | 755.36 | 100.56 |
| 60.63 | 758.30 | 100.92 |
| 60.81 | 761.23 | 101.28 |
| 60.98 | 764.15 | 101.64 |
| 61.16 | 767.05 | 101.99 |
| 61.34 | 769.95 | 102．34 |
| 61.51 | 772.83 | 102.69 |
| 4180 | アフィ 7 の | n） |

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